

Bar-Driven Dynamic Structures in Local Velocity Space

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Outline

Introduction
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Motivation
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○○

The boring stuff
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○○

Pretty Pictures and Animations
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Summary
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Outline

Introduction

Hercules stream as a testing ground

Motivation

Background

The boring stuff

Models'n'things

A quick word on methodology

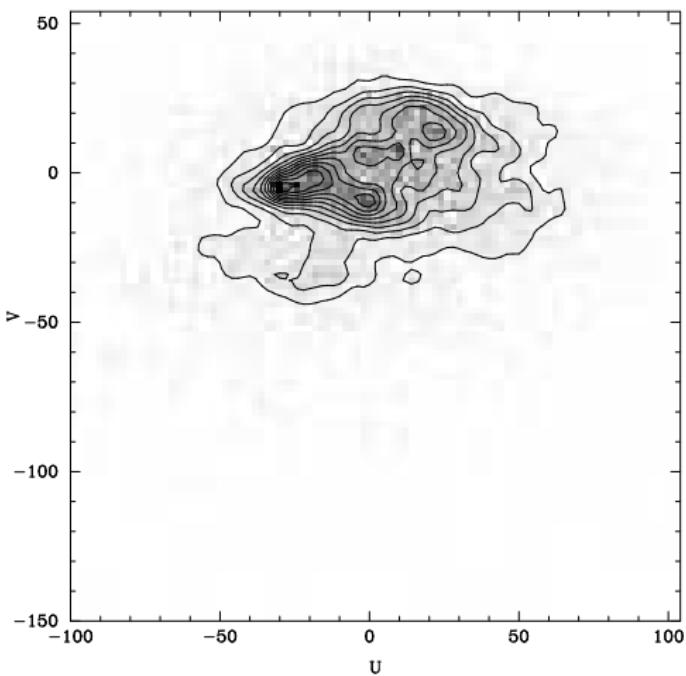
Pretty Pictures and Animations

Simulation results

Summary

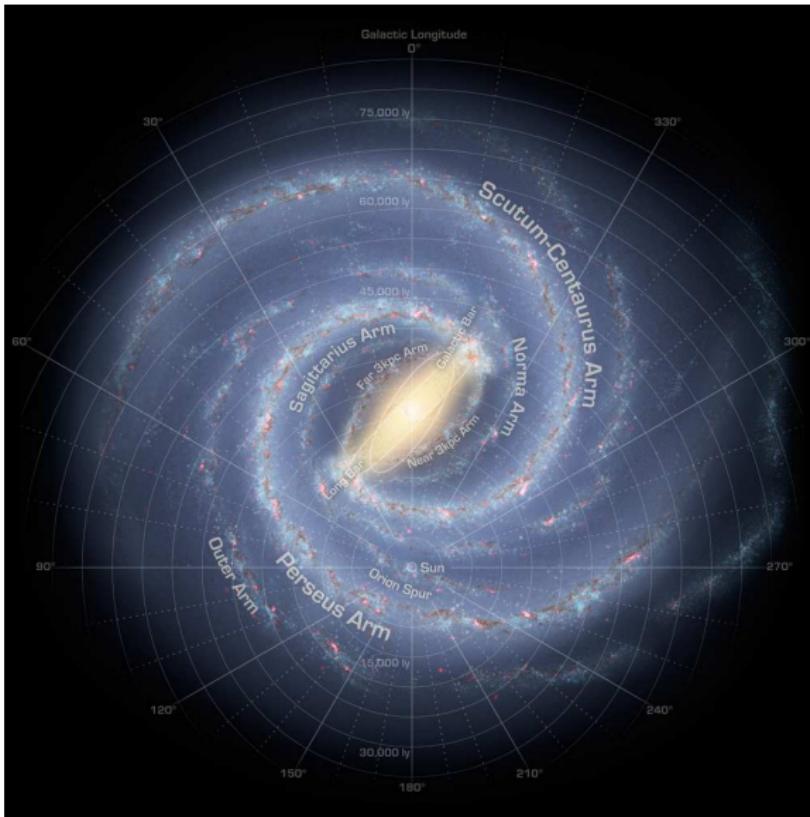
This is...

The one we're interested in

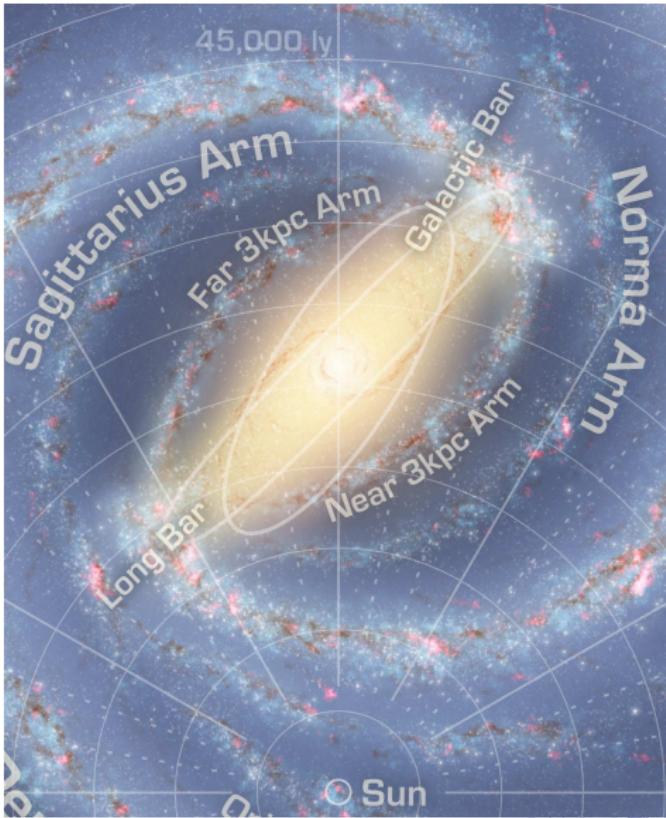


the Galaxy

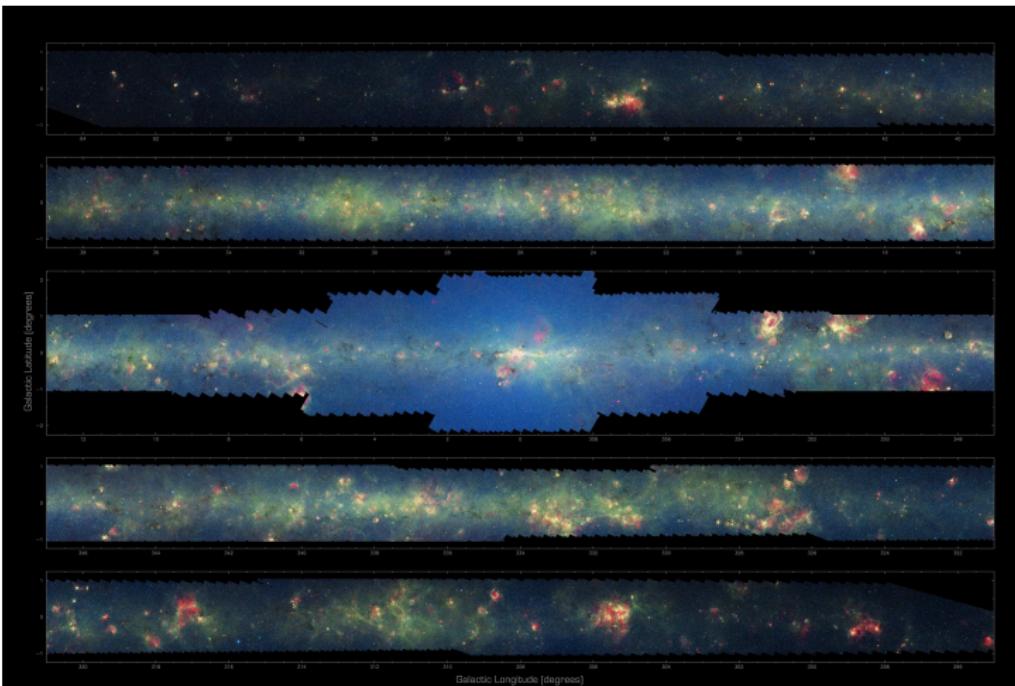
(An artist's impression of)



(An artist's impression of) the Galaxy (detail)



(What we actually see when looking at) the Galaxy



The Infrared Milky Way: GLIMPSE/MIPSGAL Spitzer Space Telescope • IRAC • MIPS

NASA / JPL-Caltech / E. Churchwell (Univ. of Wisconsin), GLIMPSE Team & S. Carey (SSC-Caltech), MIPSGAL Team ssc2008-11a

Pick a bar, any bar

- We know that a bar can cause dynamical structure
- The dynamical signature might be seen locally,
- The problem with the Milky Way seems to be that:
- There are two different observed large-scale (kpc) bars. (see e.g.: Benjamin et al. 2005, Bissantz & Gerhard 2002)
- So let's study the effects of two bars separately.
- And then go for two.

Potentials

- Standard axisymmetric stuff:
- Bulge, Halo, Dark Halo (see Flynn et al. 1996)
- Disk, scale length of 3 kpc
(new, in comparison to Flynn et al. 1996)
- Non-standard triaxial stuff:
- Ferrers' $n = 2$ potential (Pfenniger 1984, Ferrers 1877)
- $\rho = \rho_0(1 - m^2)^n$
- $m = \frac{x^2}{a^2} + \frac{y^2}{b^2} + \frac{z^2}{c^2}$
- $\rho = 0$, when $m \geq 1$

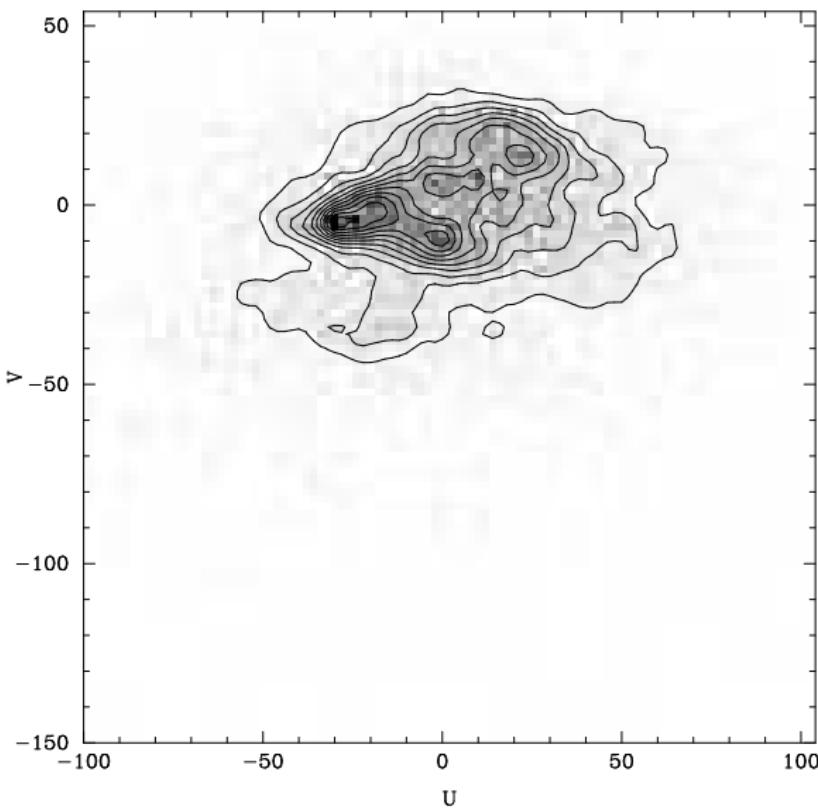
Backwards restricted method

- Dehnen (2000)
- Create a library of orbits
- Assign a weight for each point of phase-space for a certain orbit
- Compile the orbits into a table of some kind
- Interpret

Quick words on simulations

- Uniform 100×100 grid of velocity points, starting at our local position.
- Velocities $U: [-50, 50]$, $V: [-150:50]$, 2 kms^{-1} steps
- Full orbital history for each velocity point over 1 Gyr
- Weighted values for each point in velocity-space;
- Corrected for the local circular velocity, and asymmetric drift.
See Lewis and Freeman (1989).

Reruns!



Outline

Introduction
○○○○

Motivation
○
○○

The boring stuff
○
○○

Pretty Pictures and Animations
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Summary
○

What does it do? - Mass

Outline

Introduction
○○○○

Motivation
○
○○

The boring stuff
○
○○

Pretty Pictures and Animations
○○●○○○○○○○○

Summary
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OLR variations - Long Bar

Outline

Introduction
○○○○

Motivation
○
○○

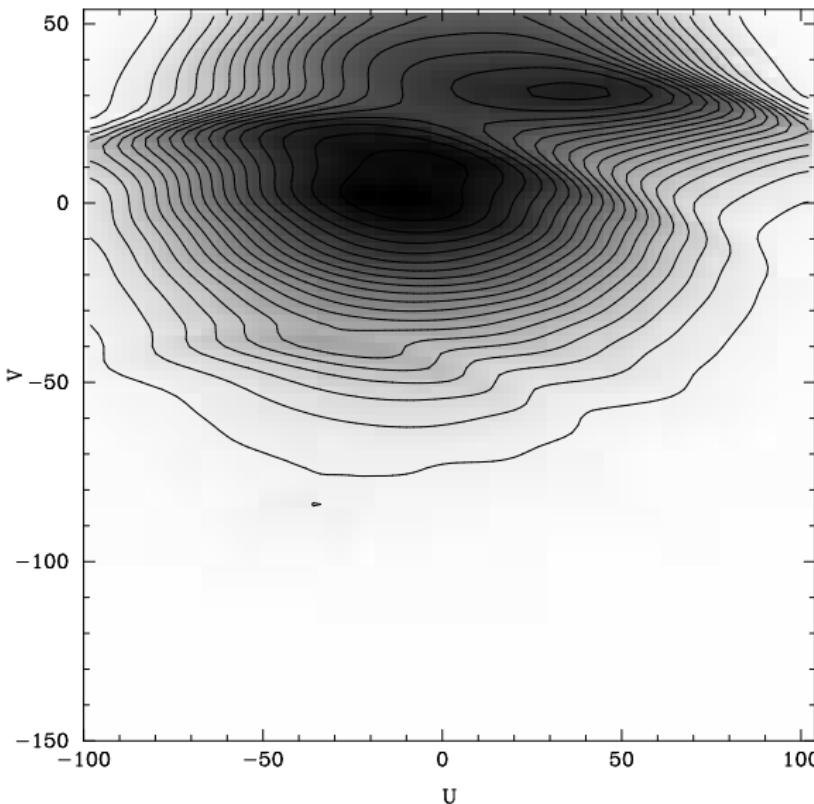
The boring stuff
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Pretty Pictures and Animations
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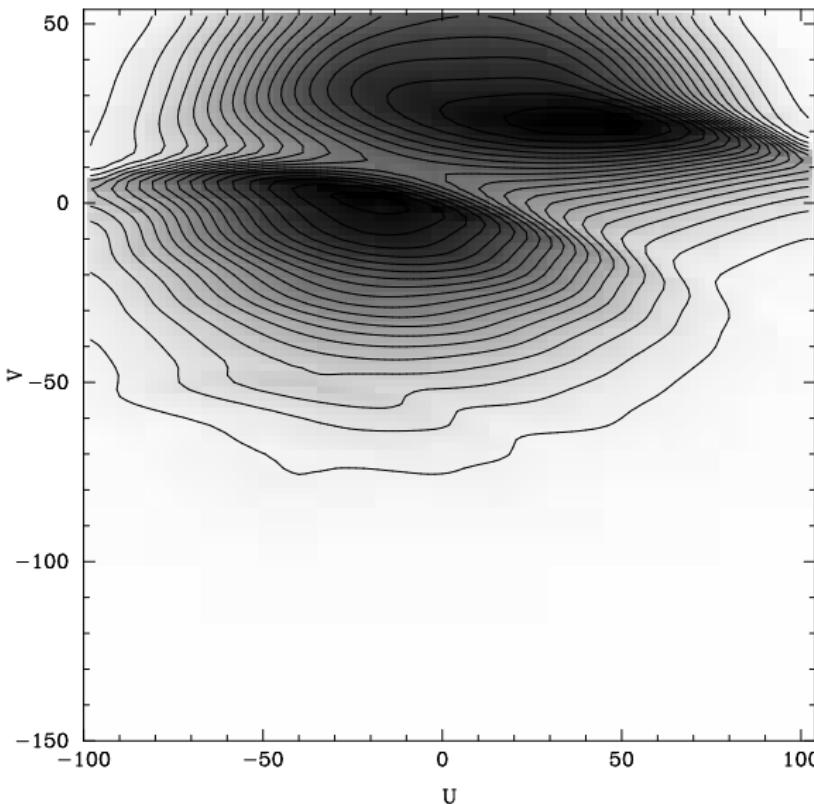
Summary
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Angle variations - Long Bar

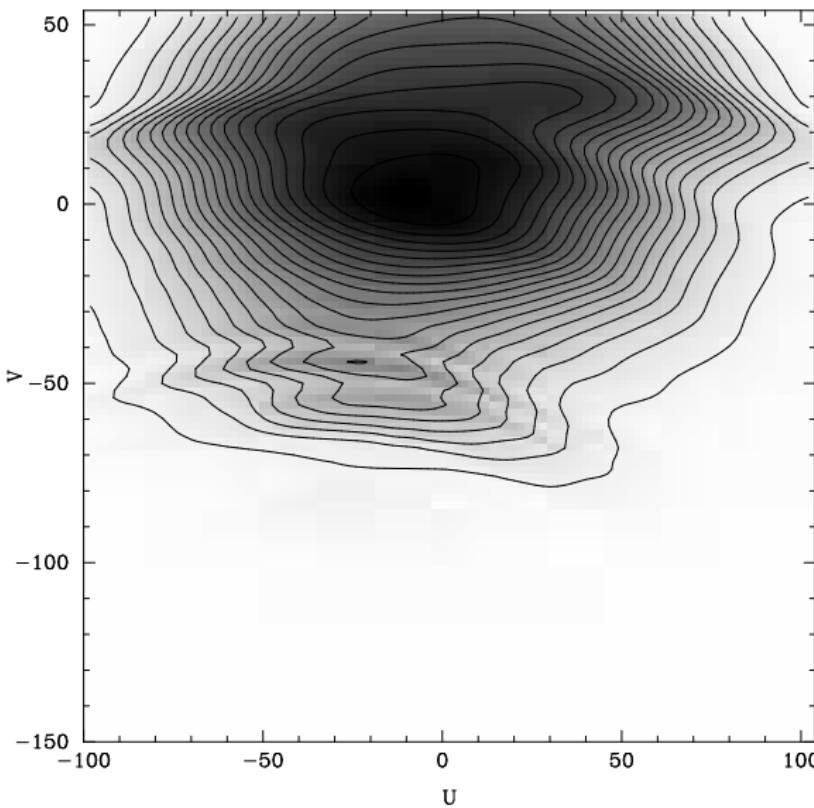
Two bars! Phase locked at 1.87.



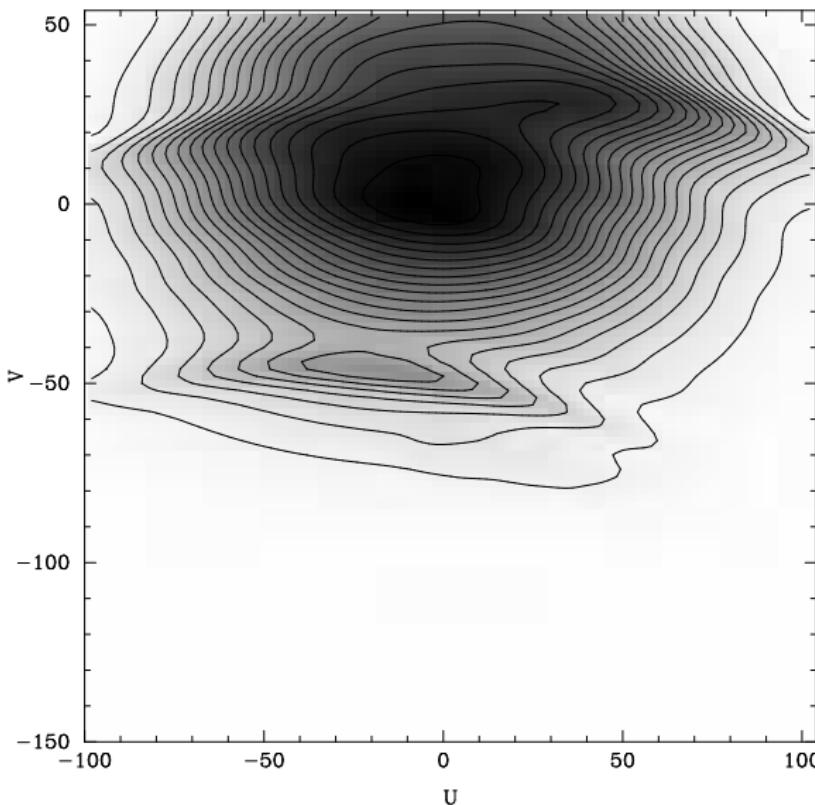
Two bars! Phase locked at 1.95.



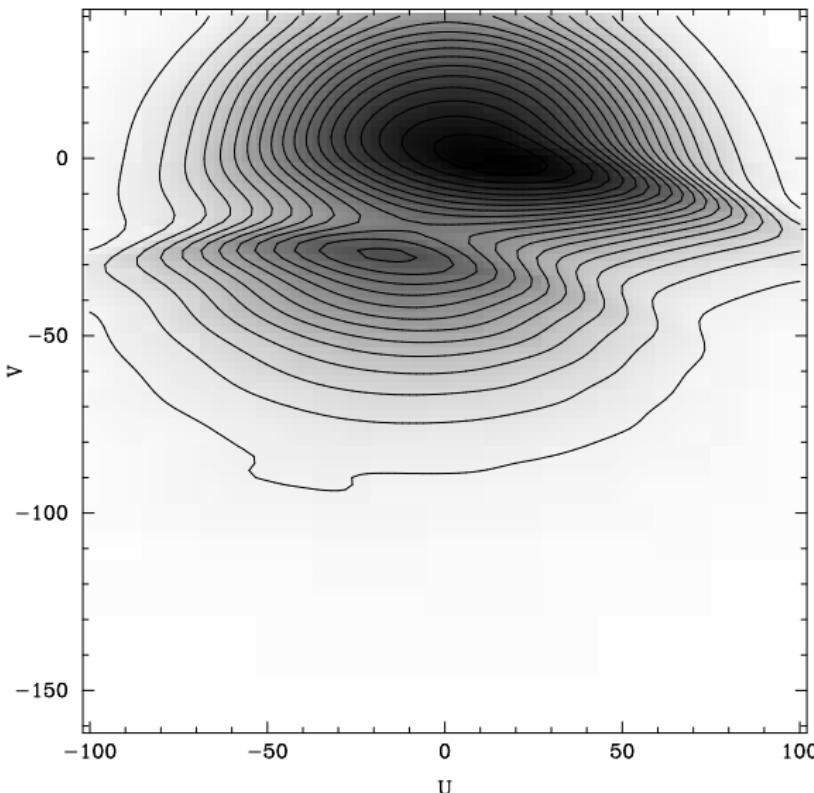
Two bars! Phase unlocked, long at 1.87, Galactic at 1.50



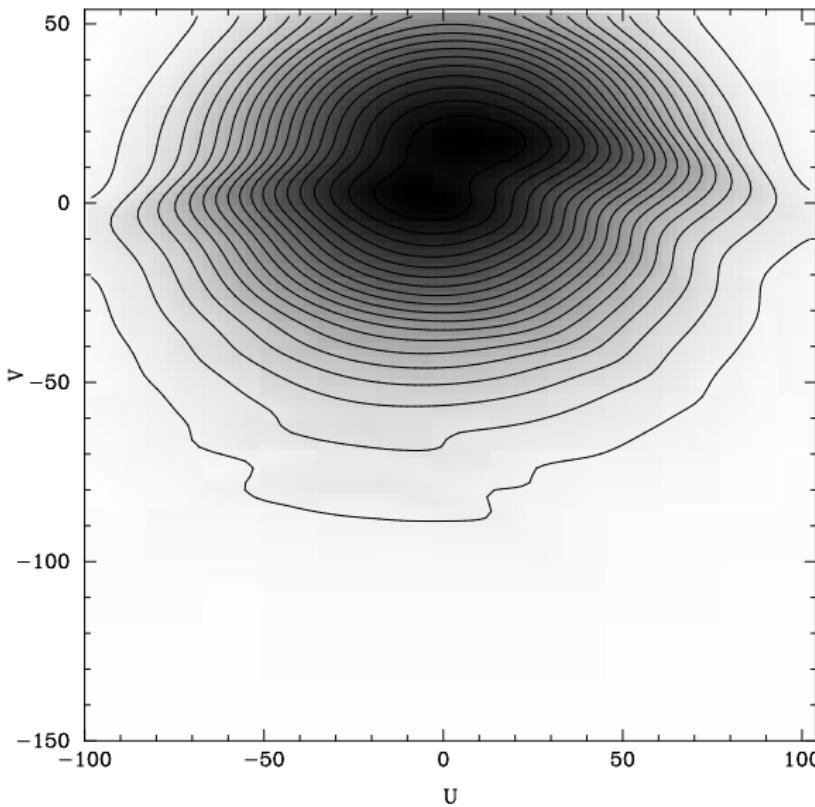
Two bars! Phase unlocked, long at 1.50, Galactic at 1.87



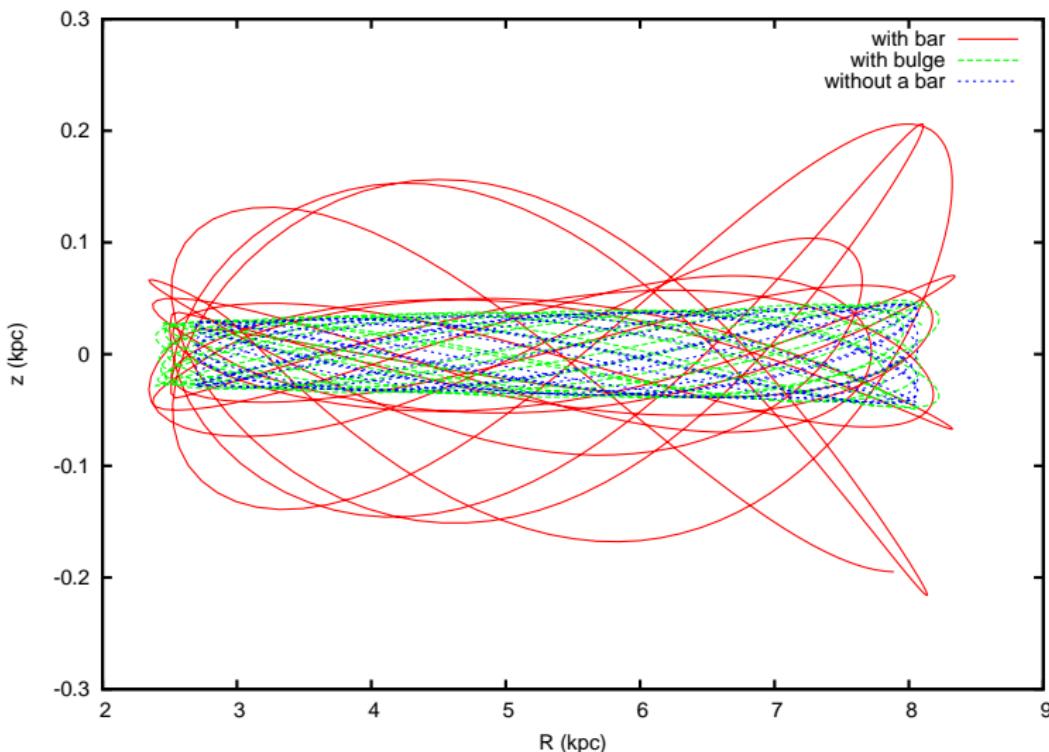
Two bars! Phase locked at 1.87, half-mass.



Two bars! long at 1.87, Galactic at 1.50, half-mass.



Potentially something very interesting



The End

- A bar will create structure(s) in velocity space.
- The angle, mass, and speed of the bar will affect it's position and shape in velocity space.
- Two large, massive bars will wreck havok on velocity structure.
- Direct interaction with bars can cause also cause structure in velocity space.
- More details in Gardner & Flynn 2010 (MNRAS 405, p. 545 or arXiv:1002.0551)