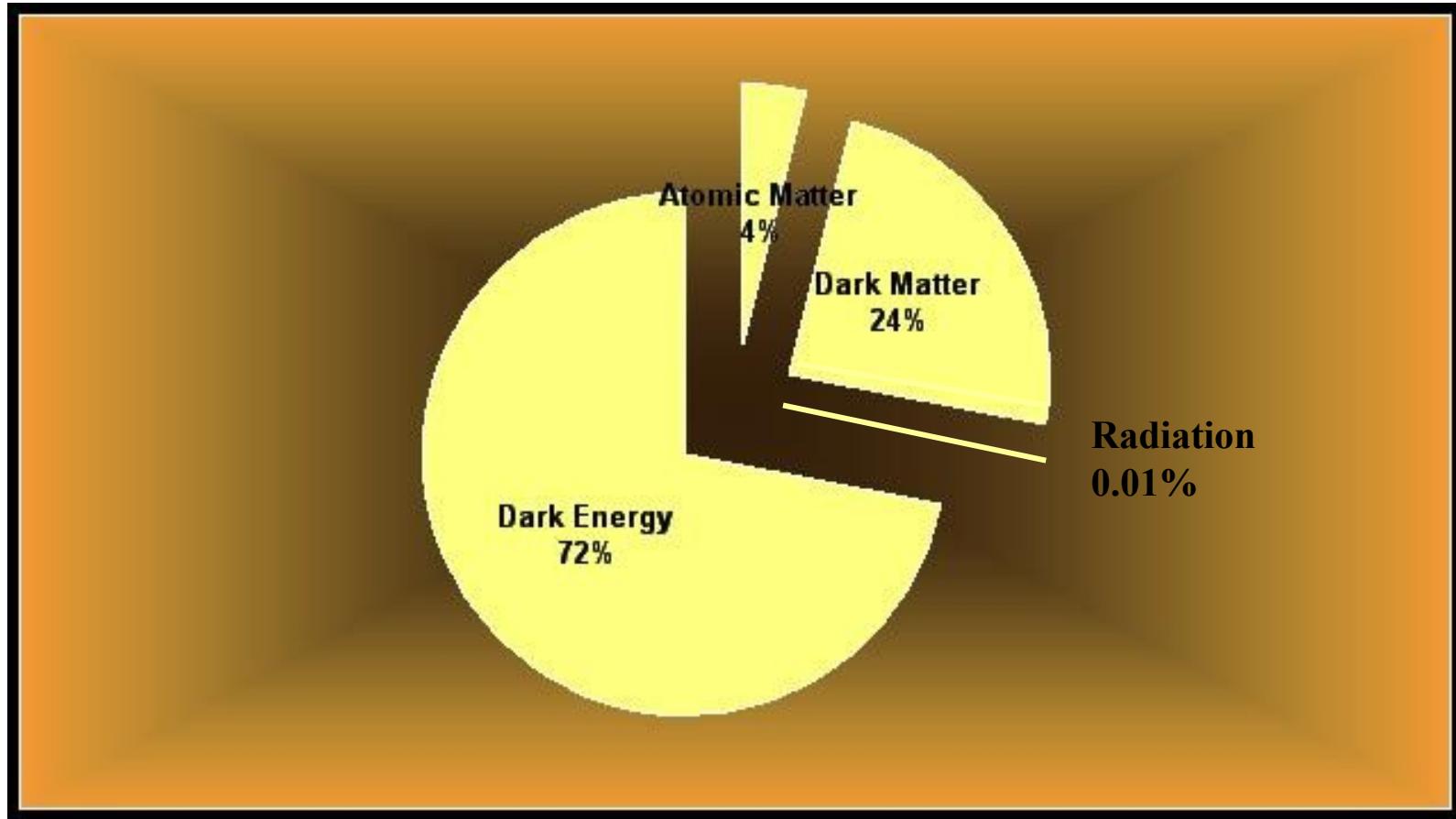




# Энергетический состав Вселенной: новая симметрия

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# Energy composition of the observed Universe



Accidental? Preposterously complex? Absurd?

In the past, at the Big Bang Nucleosynthesis  
 $(t_{\text{BBN}} \sim 100 \text{ sec})$ :

$$\rho_R \sim \rho_{\text{TOT}}$$

$$\rho_{\text{DE}} \sim 10^{-32} \rho_{\text{TOT}}$$

$$\rho_{\text{DM}} \sim \rho_B \sim 10^{-6} \rho_{\text{TOT}}$$

In future, at, say,  $t \sim 100 t_0$ :

$$\rho_R \sim 10^{-116} \rho_{\text{TOT}}$$

$$\rho_{\text{DE}} \sim \rho_{\text{TOT}}$$

$$\rho_{\text{DM}} \sim \rho_B \sim 10^{-90} \rho_{\text{TOT}}$$

*Are there any time-independent energy parameters?*

*Yes, these are the FRIEDMANN INTEGRALS*



## Friedmann equation (I)

$$K + U = 0$$

$$(dR/dt)^2 = (A_{DE}/R)^{-2} + A_{DM}/R + A_B/R + (A_R/R)^2$$

$$DE = \Lambda, \quad \Omega = 1, \quad c = 1$$

# The Friedmann integrals

$$A_{DE} = M_P (\rho_{DE})^{-1/2}$$

$$A_{DM} = (3/4\pi) M_P^{-2} M_{DM}$$

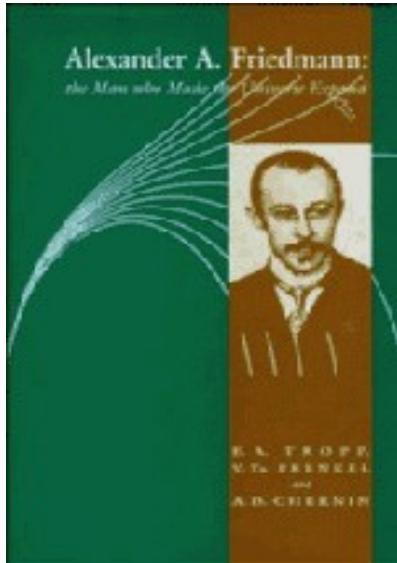
$$A_B = (3/4\pi) M_P^{-2} M_B$$

$$A_R \approx (3/4\pi) M_P^{-1} N_R^{2/3}$$

$M_{DM}$ ,  $M_B$ ,  $N_R$  are global (say, for the whole observed universe)

covariant conserving values

Natural units:  $c = \hbar = 1$ ;  $M_P = (8\pi G/3)^{-1/2} \sim 10^{18} \text{ GeV}$



## Friedmann equation (II)

$$dS = 0$$

$$R \frac{d\rho}{dt} = -3 \rho (1+w) \frac{dR}{dt};$$

Solution: the Friedmann integral

$$A = [M_P^{-2} \rho R^{3(1+w)}]^{1/(1+3w)}$$

Equation of state:  $w = p/\rho$

DE=Λ:  $w = -1$ ; DM:  $w = 0$ ; B:  $w = 0$ ; R:  $w = 1/3$

Normalized scale factor:  $R(z) = 3/H_0 (1+z)^{-1}$

# Time-independent empirical recipe of the cosmic energy mix

$$A_{DE} \sim A_{DM} \sim A_B \sim A_R \sim 10^{60 \pm 1} M_P^{-1}$$

**Internal (non-geometrical) symmetry of the cosmic energies**

**Dark energy is a regular member of the cosmic energy quartet**

## A special feature of the present epoch:

The four energy densities are approximately equal,

$$\rho_{\text{DE}} \sim \rho_{\text{DM}} \sim \rho_B \sim \rho_R ,$$

when the Hubble radius is near the Friedmann integral,

$$R_H \sim 1/H_0 \sim A_{\text{DE}}$$

# Link to fundamental physics

Okun (1988), Weinberg (2000):

Central part of electroweak energy scale  $M_{EW} \sim 1000 \text{ GeV}$

Arkani-Hamed et al. (2000):  $M_{EW}$  in “dark physics”

WIMP mass  $m_{DM} \sim M_{EW}$

DE density  $\rho_{DE} \sim X^{-8} M_P^{-4} \sim 10^{-120} M_P^{-4}$

Hierarchy number  $X \sim M_P/M_{EW} \sim 10^{15}$

$$1) \quad A_{DE} \sim X^4 M_P^{-1}$$

2) Dark matter freeze-out at  $z \sim X$

$$A_{DM} \sim A_R \sim X^4 M_P^{-1}$$

1) + 2)

$$A_{DM} \sim A_R \sim A_{DE} \sim X^4 M_P^{-1} \sim 10^{60} M_P^{-1}$$

3) Baryons: electroweak baryogenesis??

# Cosmology big numbers & Hierarchy phenomenon

Big Baryonic Number:  $B = n_R/n_B \sim (m_B/M_P) X \sim 10^{12}$

Big Dark Number:  $D = n_R/n_{DM} \sim X \sim 10^{15}$

Total number of relic photons:  $N_R \sim X^6 \sim 10^{90}$

Total number of WIMPs:  $N_{DM} \sim X^5 \sim 10^{75}$

Hierarchy number  $X \sim M_P/M_{EW} \sim 10^{15}$

## Cosmic internal symmetry

$$A_{DE} \sim A_{DM} \sim A_B \sim A_R \sim (M_P/M_{EW})^4 M_P^{-1} \sim 10^{60} M_P^{-1}$$

- \* Unifies the four cosmic energies
- \* Exists whenever the four energies exist
- \* Covariant
- \* Robust
- \* Approximate
- \* Originated from gravity-electroweak interplay at the first few picoseconds and related to the hierarchy phenomenon?

# References

- Chernin A.D. *Nature* **270**, 250, 1968 ( $A_B \sim A_R$ )
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